

H.A.L. Astronomy School



The Ones That Move:

One Current Asteroid and Two
Current Planets

Asteroid Vesta:

Nearly the Size of Arizona and
Currently Passing Near the
Denebola/Vindemiatrix Line





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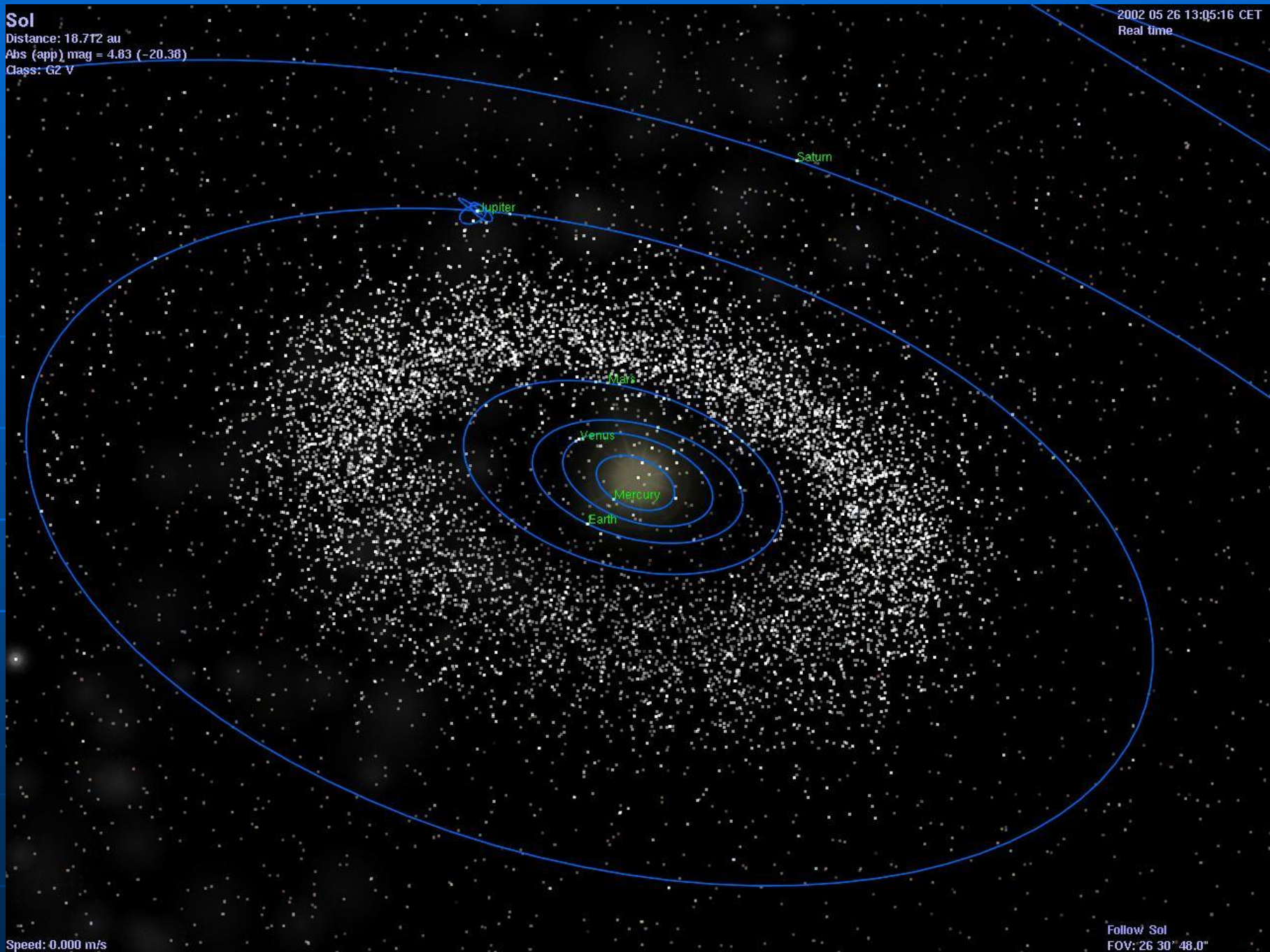
REPRODUCED WITH PERMISSION NO. 27280



Sol

Distance: 18.712 au
Abs (app) mag = 4.83 (-20.38)
Class: G2 V

2002 05 26 13:05:16 CET
Real time

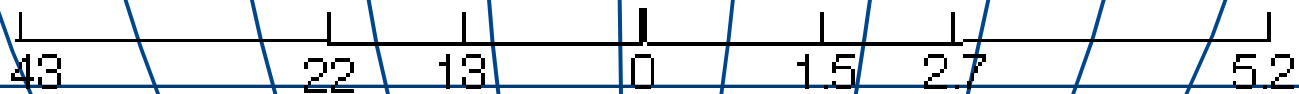
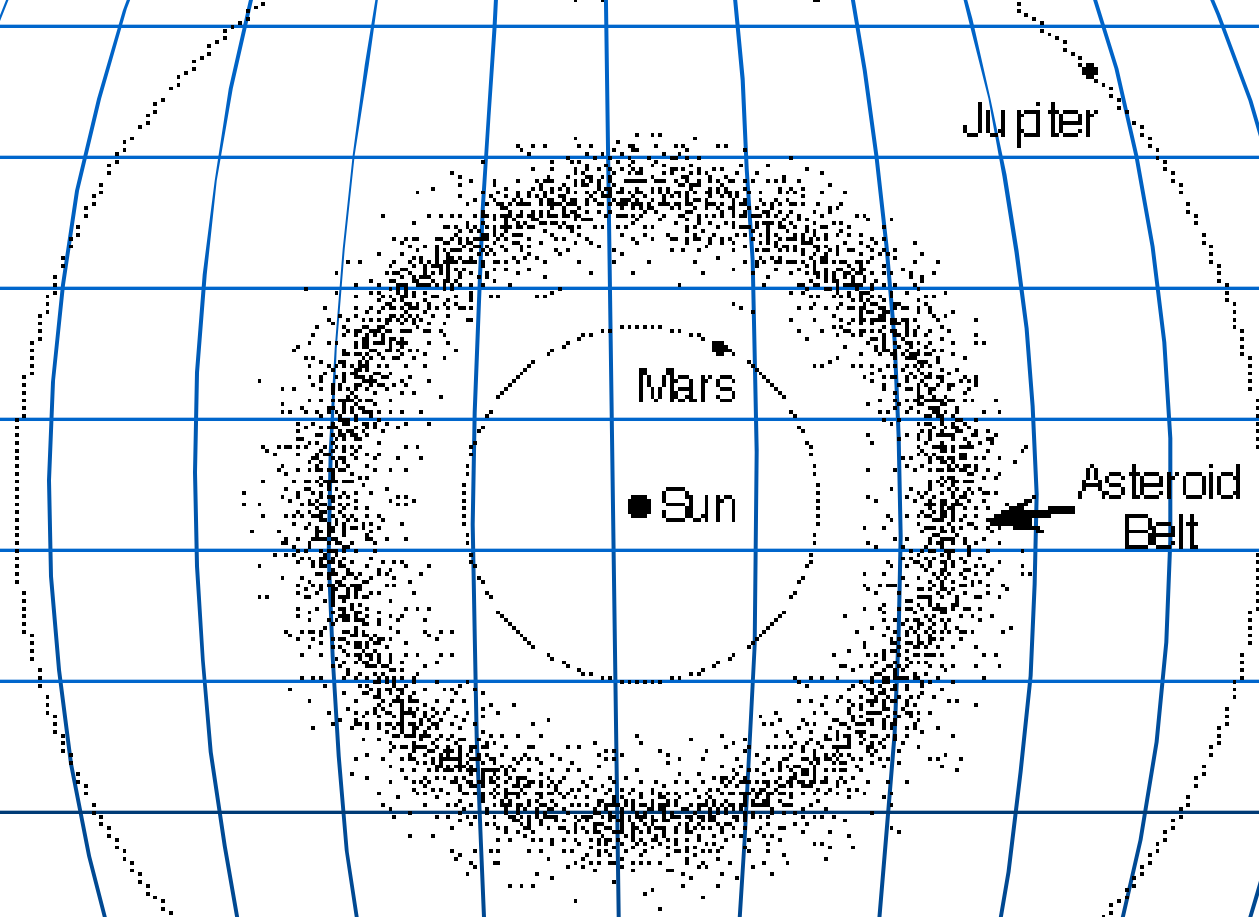


Speed: 0.000 m/s

Follow Sol
FOV: 26 30" 48.0"

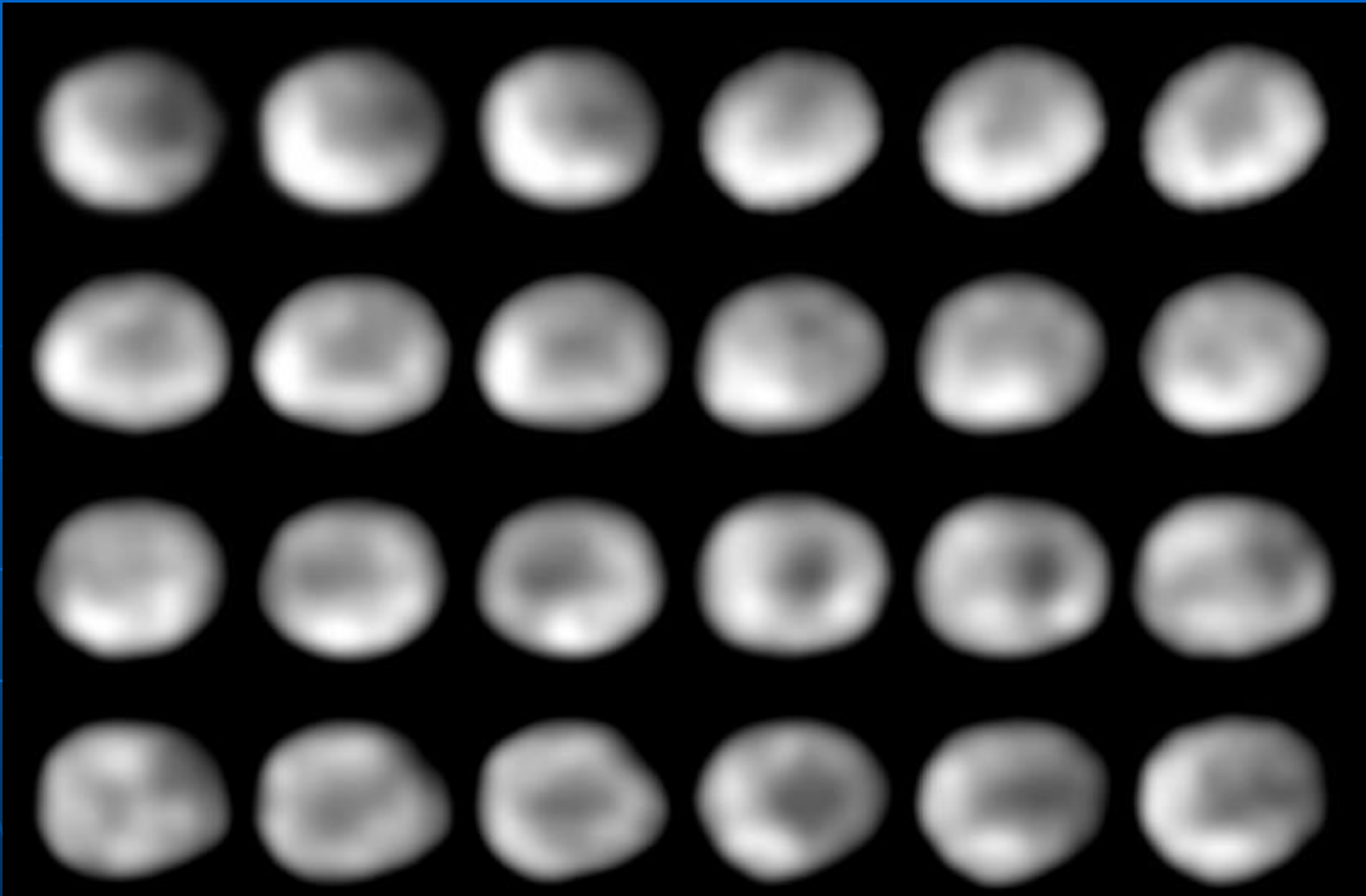
The Main Asteroid Belt

(Orbits drawn approximately to scale)



Light Minutes

Astronomical Units



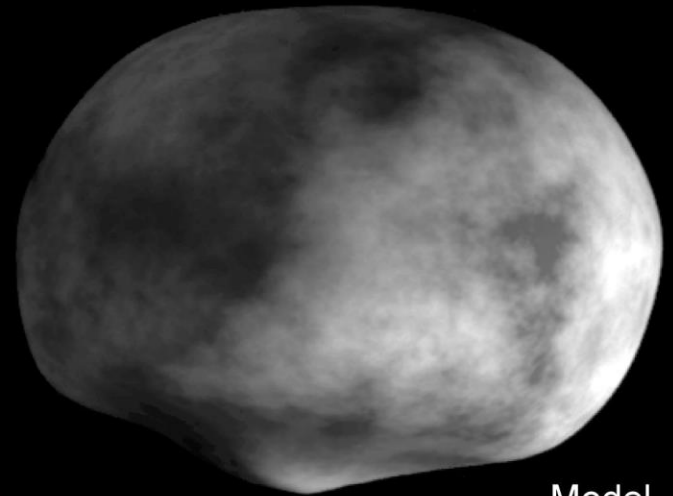
Asteroid Vesta

HST · WFPC2

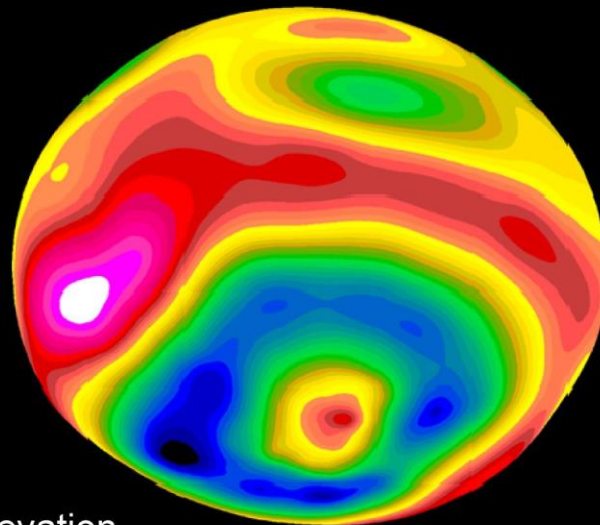
PRC95-20A · ST ScI OPO · April 19, 1995 · B. Zellner (GA Southern Univ.), NASA



HST



Model

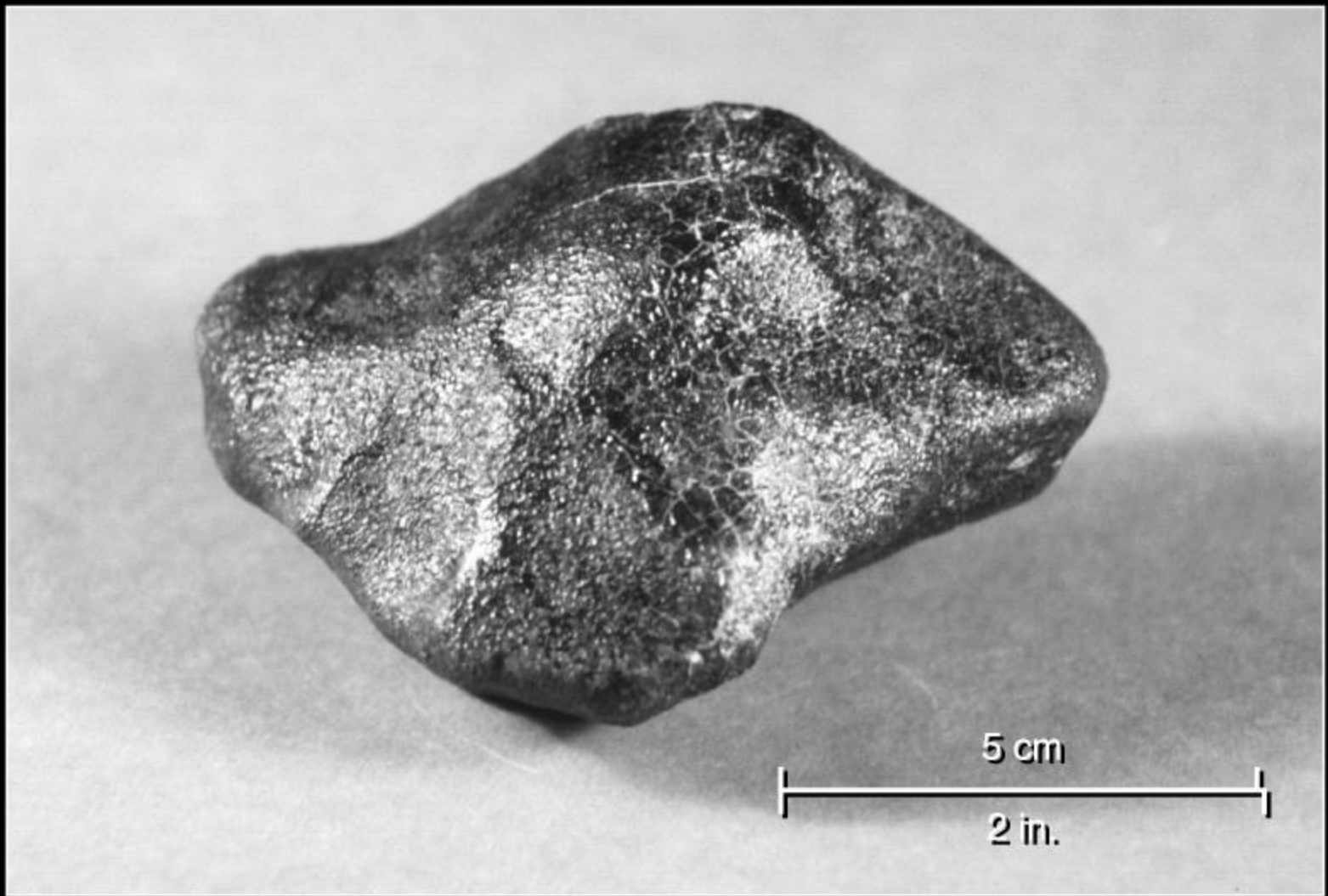


Elevation



Asteroid Vesta

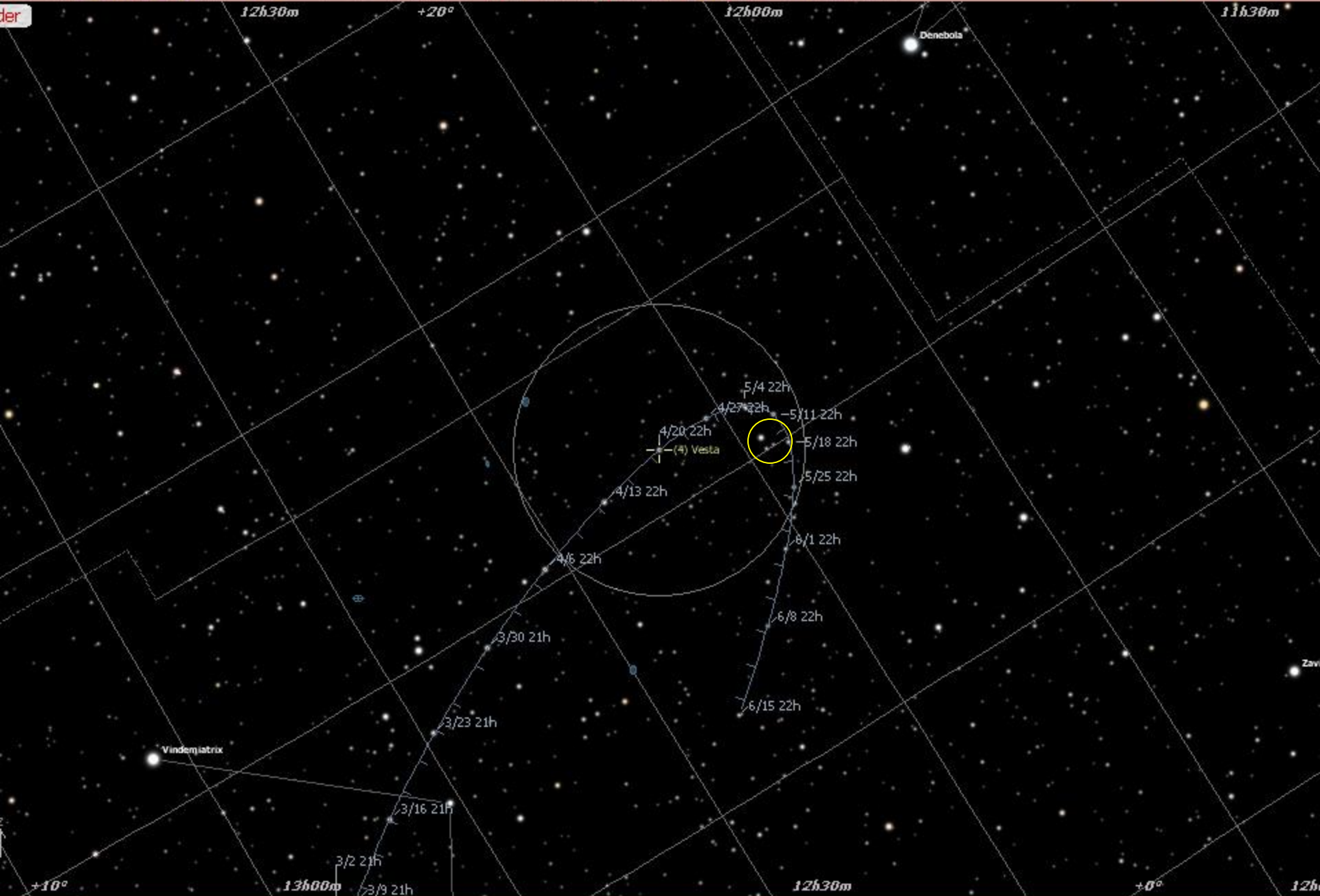
Hubble Space Telescope • Wide Field Planetary Camera 2

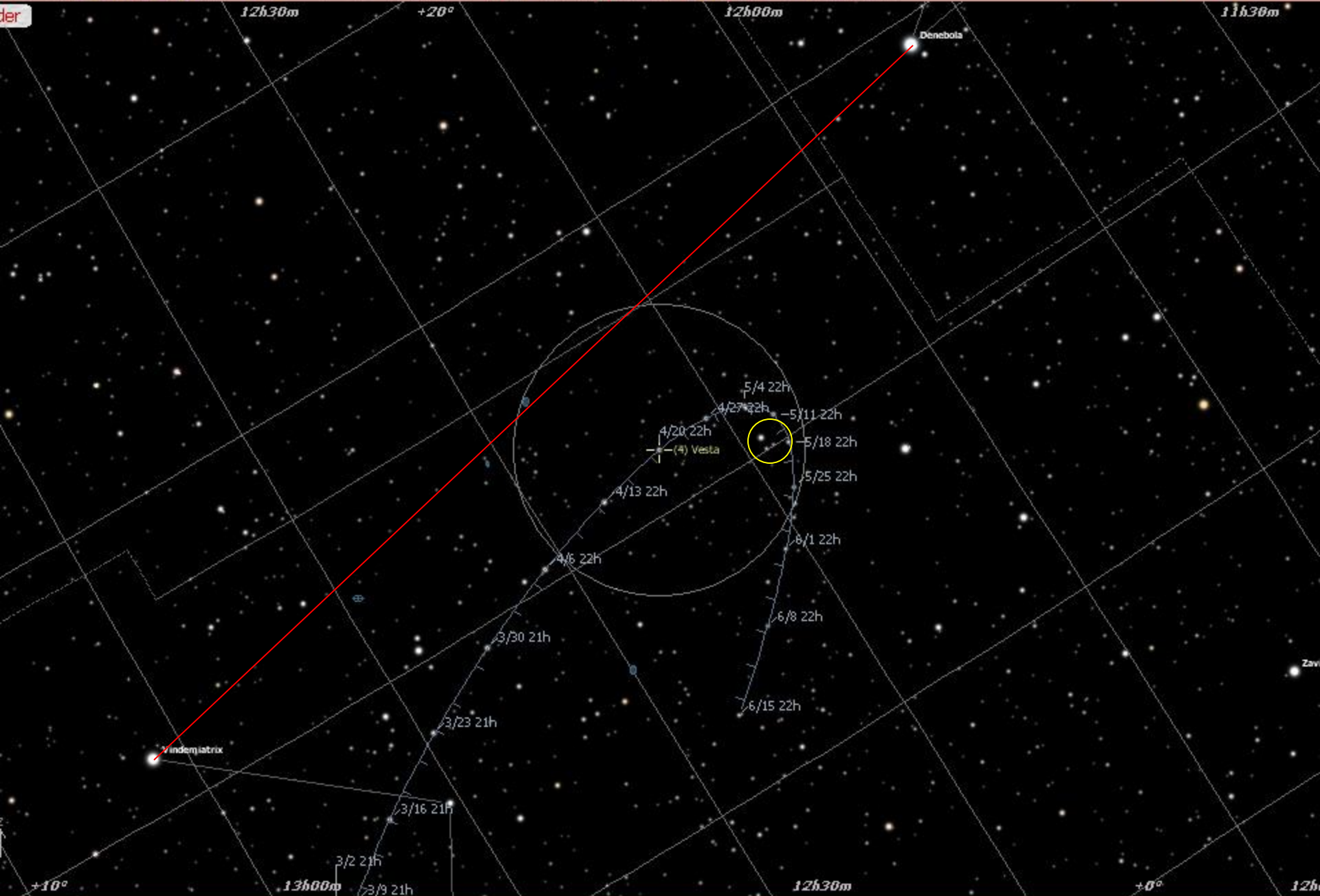


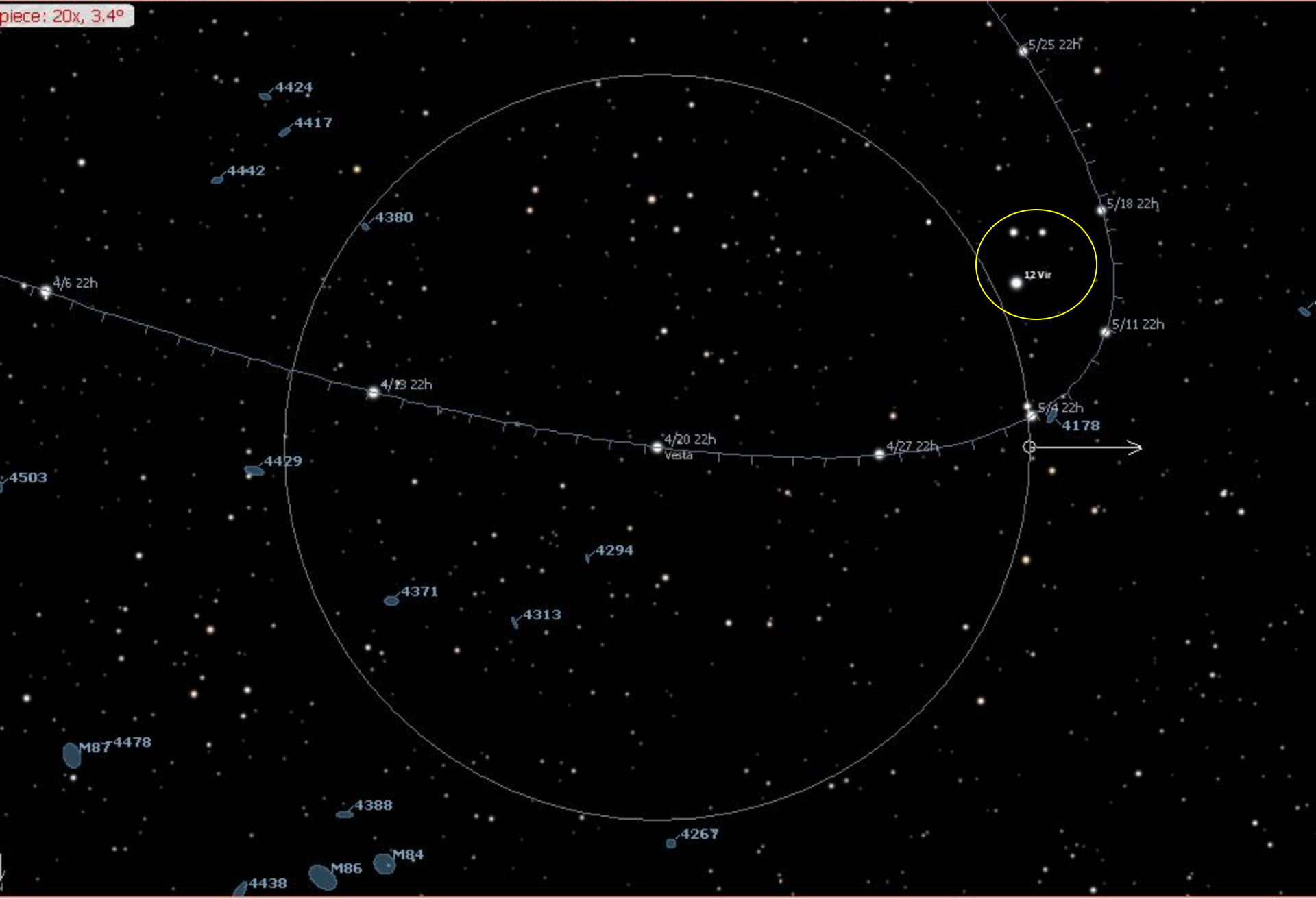
Meteroite · Fragment of Vesta

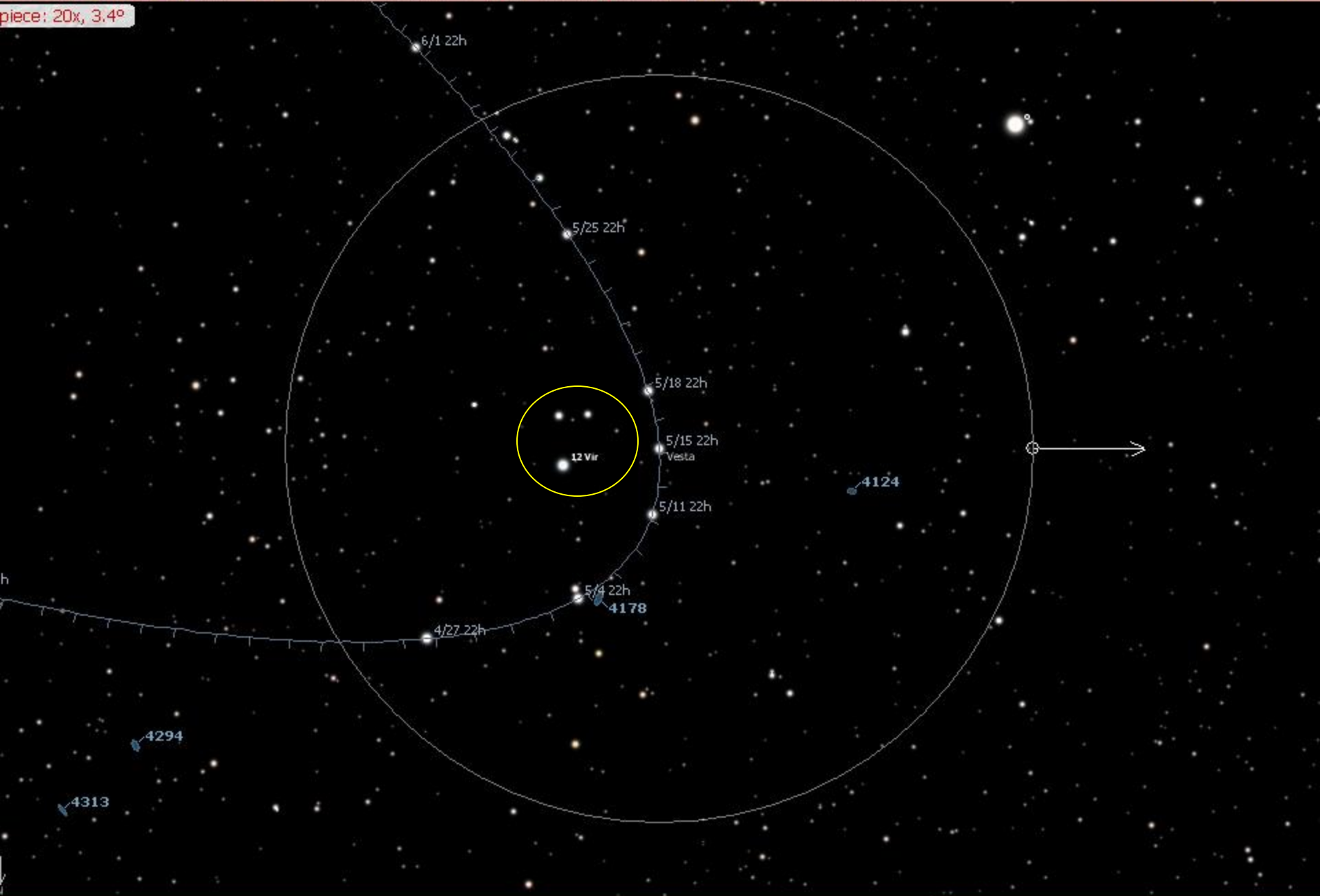
Lab Photograph · Russel Kempton, New England Meteoritical Services

PRC95-20B · ST Scl OPO · April 19, 1995 · B. Zellner (GA Southern Univ.), NASA









piece: 20x, 3.4°



ASTEROIDS: Course Plotting

Finding and following one of the small rocky planetoids that accompany the major planets around the Sun can be a most satisfying project.

The small size of asteroids can make them a challenge to find, however. Although the largest, **Ceres**, is about 1000 kilometers (620 miles) in diameter, most range from about 100 kilometers (62 miles) to 200 k (125 miles) across, down to one kilometer (0.6 mile) or less. This means they all remain starlike even in the largest of amateur scopes.

The four largest can be found in binoculars under dark skies specially at opposition when they are the brightest. All four are magnitude 8.5 or brighter.

Since they are stellar in appearance their true nature can only be discerned by their movement compared to the background stars from one night to the next.

Each year the daily or weekly positions for these fascinating little worlds are published in the astronomical periodicals.

Using the information thus obtained, find and track an asteroid over a period of 3-5 nights.

ASTEROIDS: Measuring their Movement

Having plotted an asteroid's pathway among the background stars for at least three evenings you can now figure out its approximate hourly movement.

Using a finely graded ruler such as a millimeter rule, measure the distance from the dot representing your first observation to the dot representing your second observation.

If these two observations were, for example, about 24 hours apart, divide that measurement by 24 (or whatever the time interval was in hours) to find out how far the asteroid traveled in one hour.

Do the same thing for each subsequent observation.

How far did the asteroid move in one hour?

Pluto:

Your Summer Opportunity to
Join the Nine Planets Club

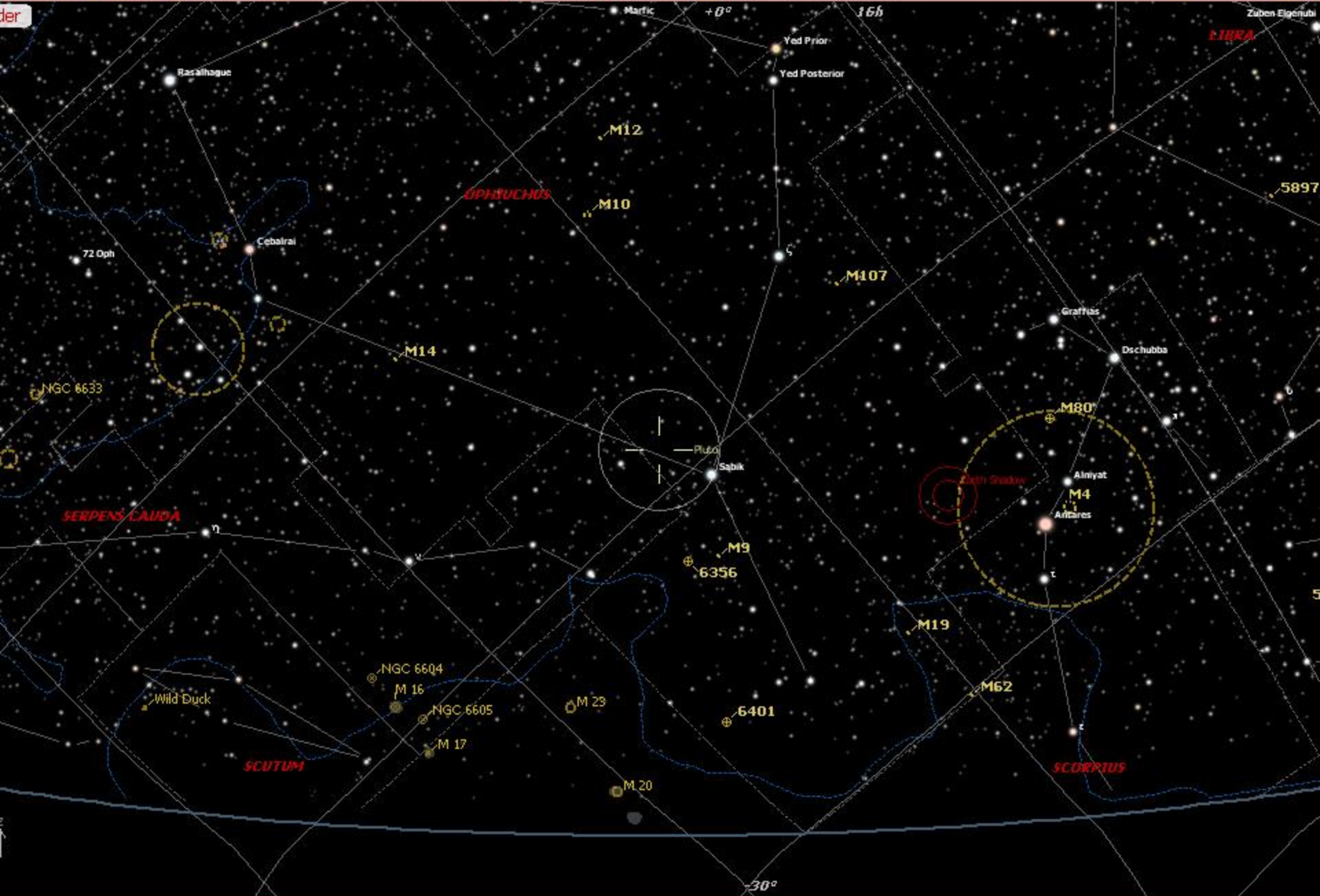
Pluto

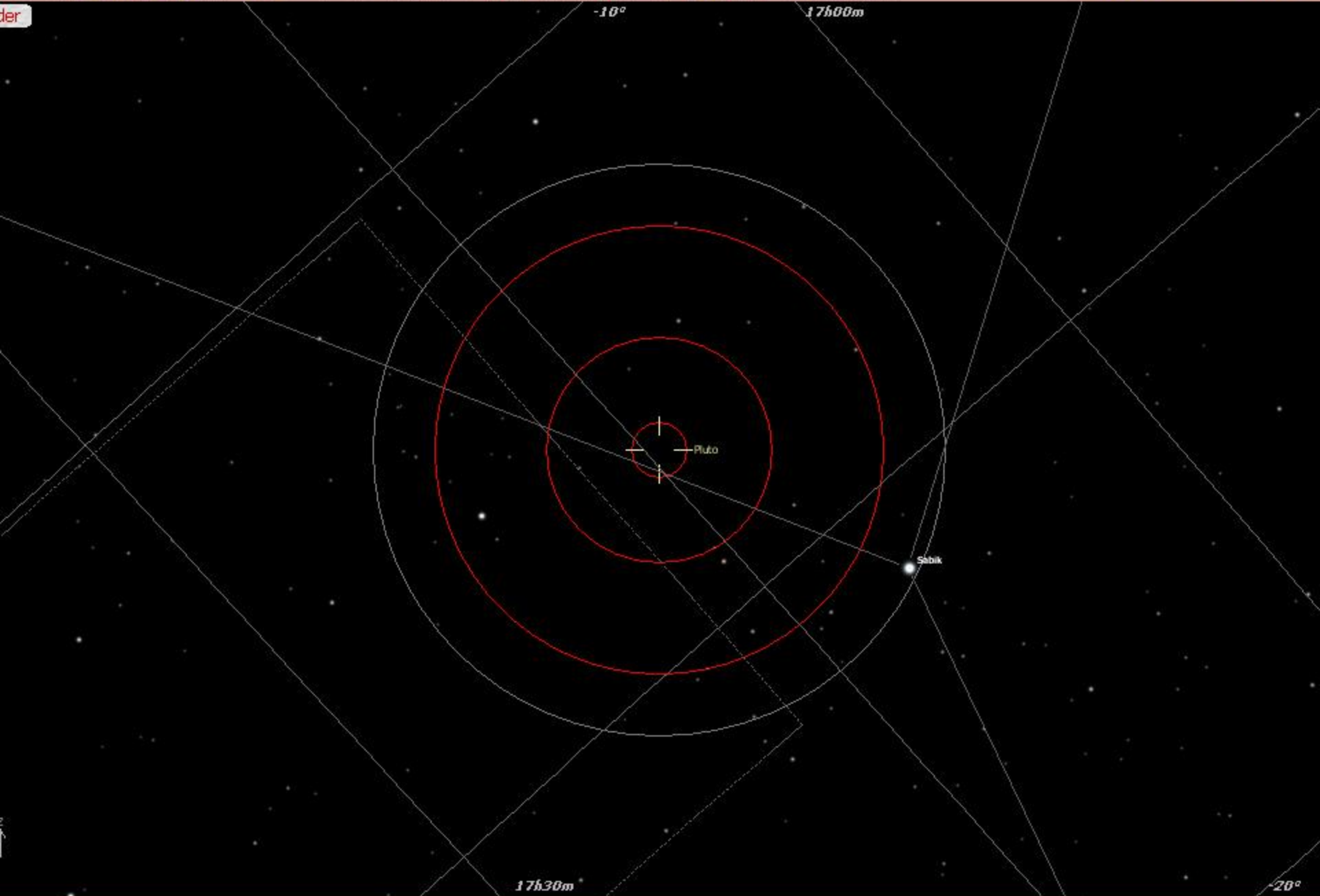
Distance: 5685.000 km

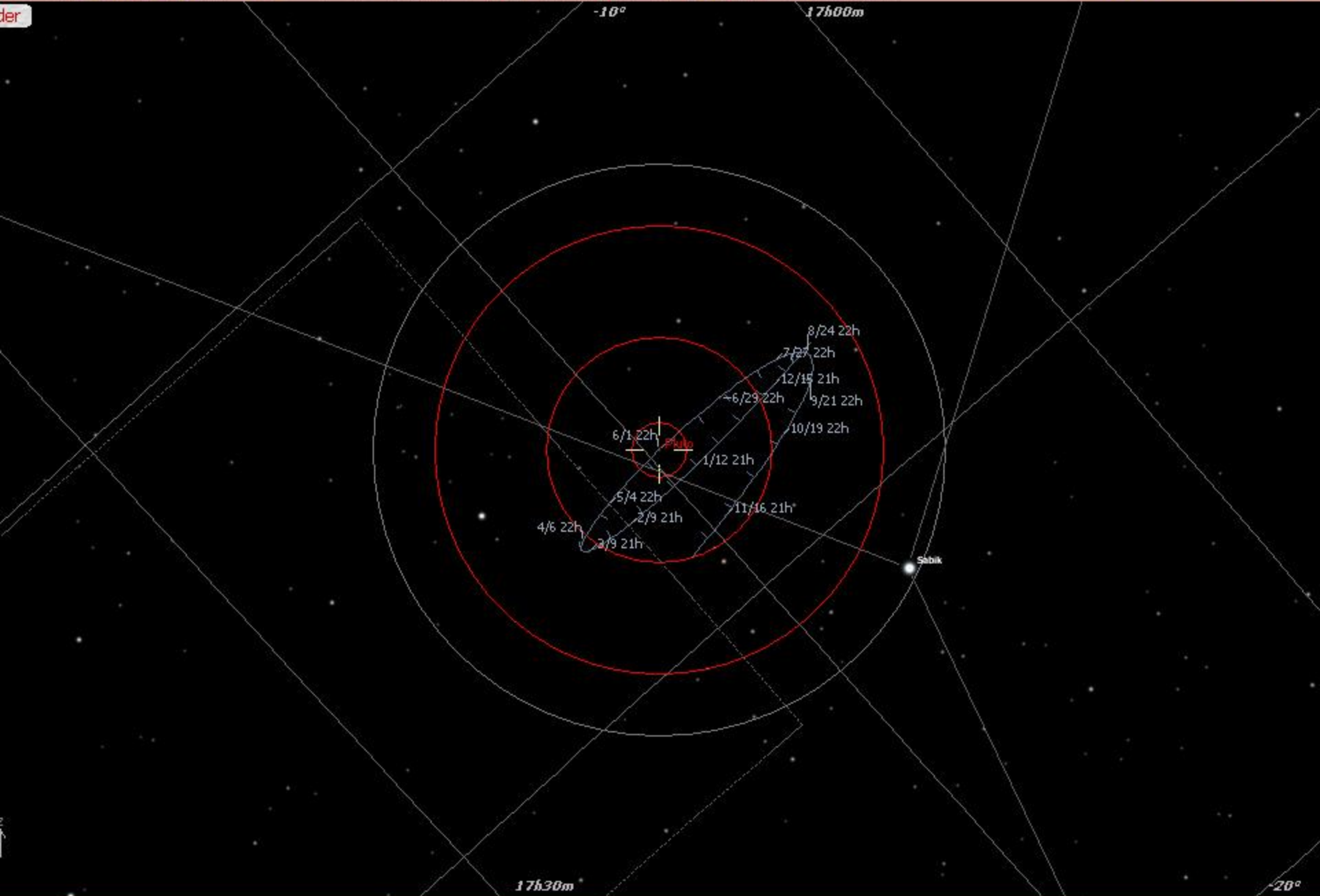
Radius: 1137.000 km

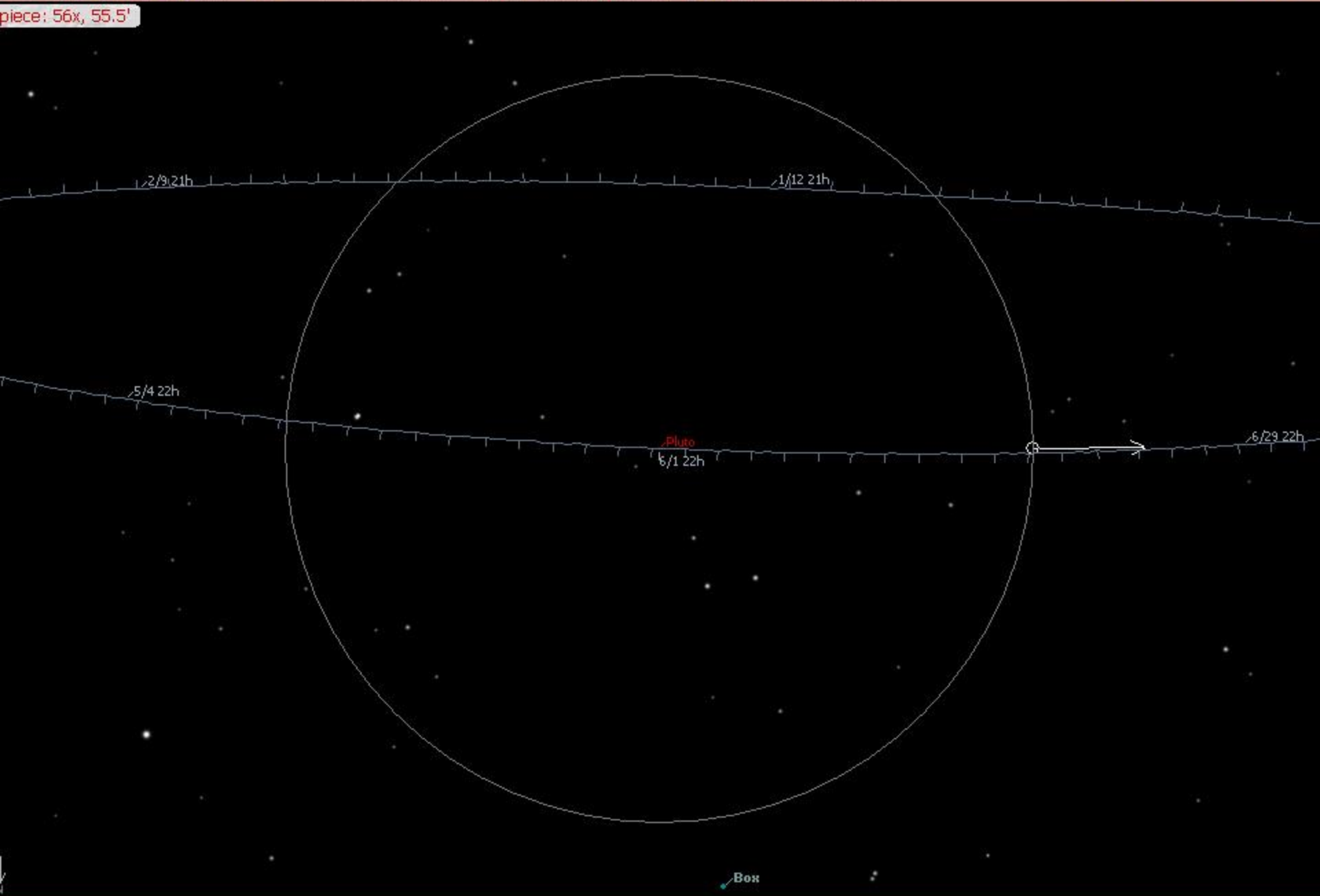
2002 05 21 09:43:28 (-
100.000x faster



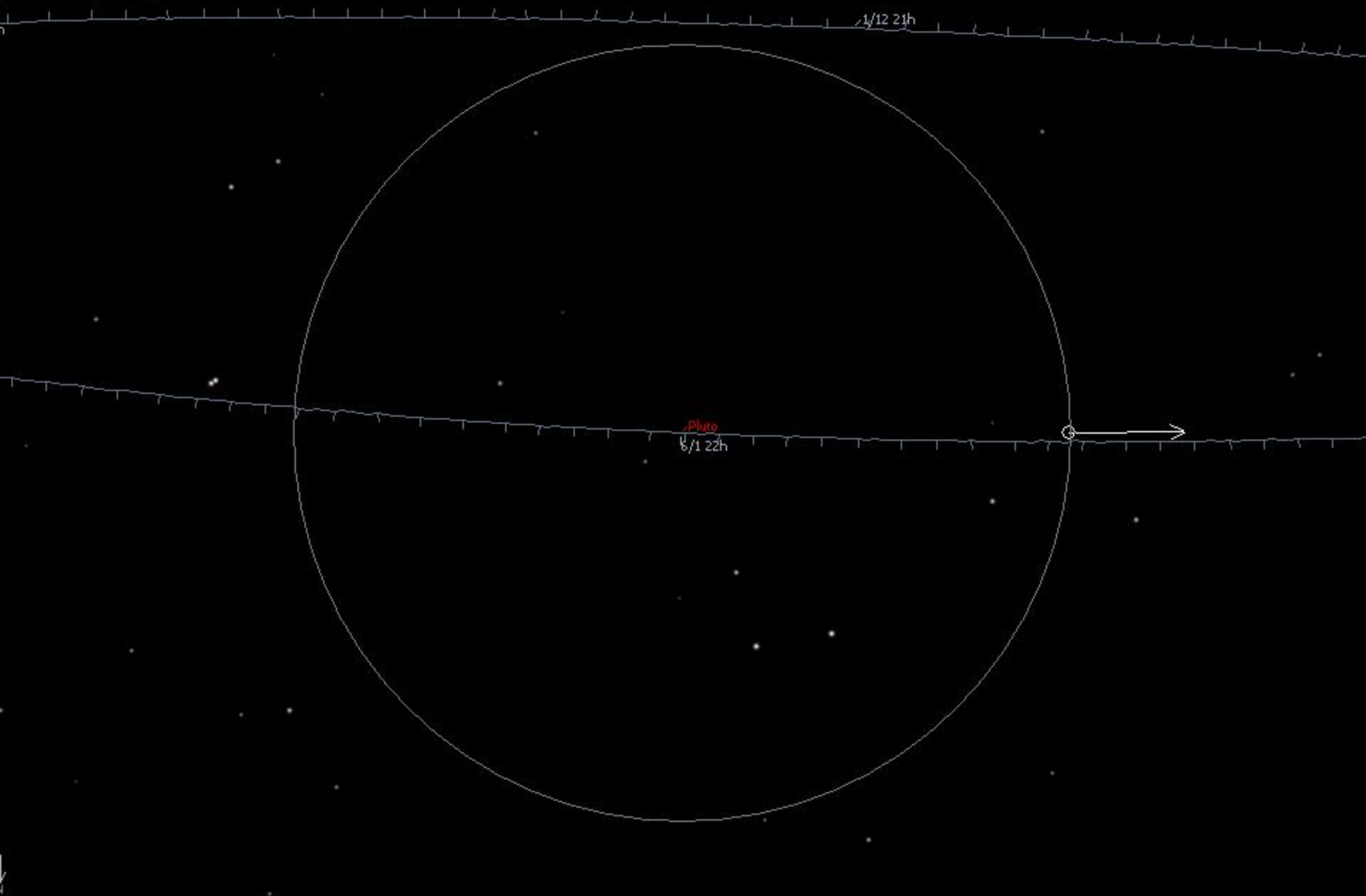






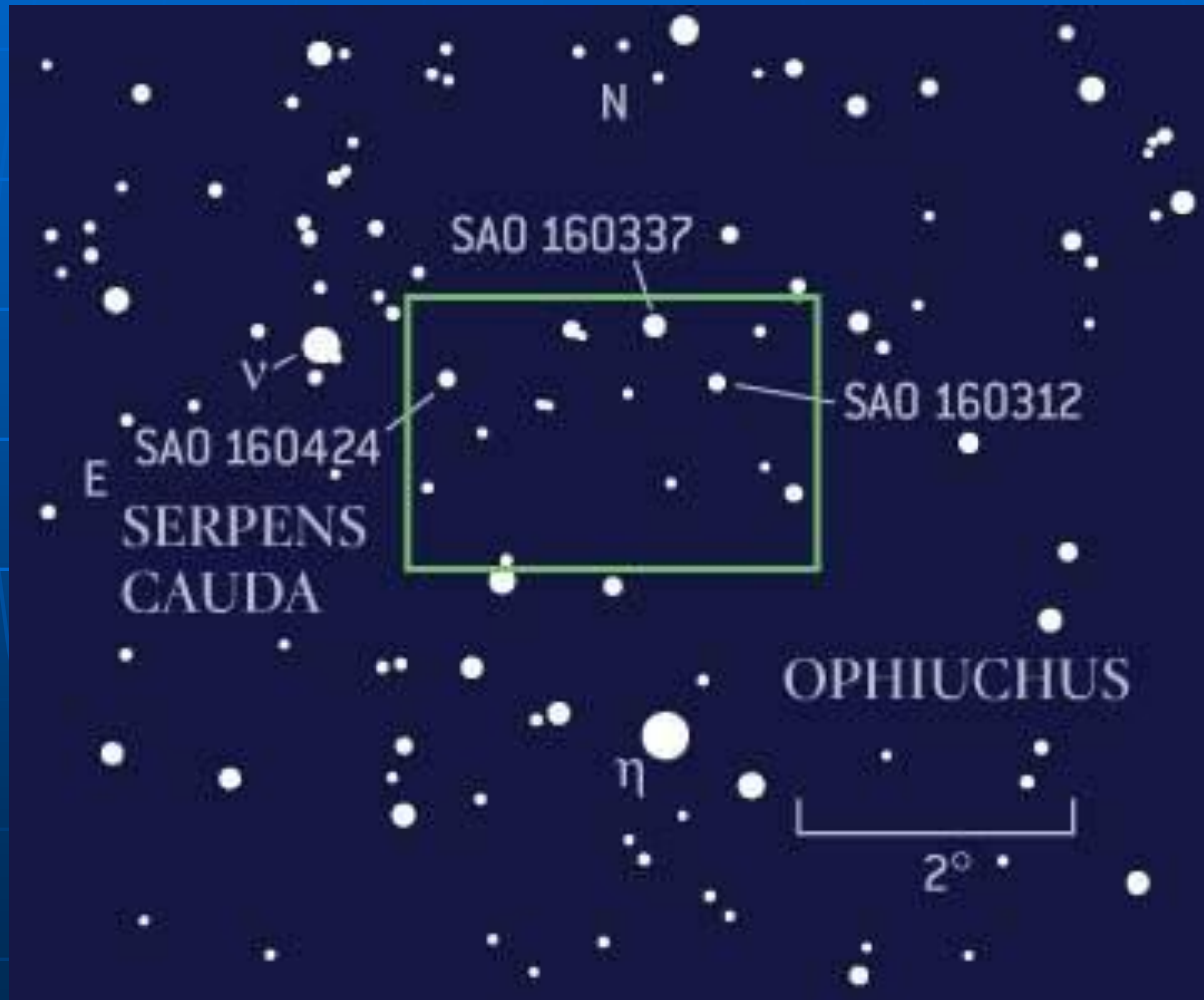


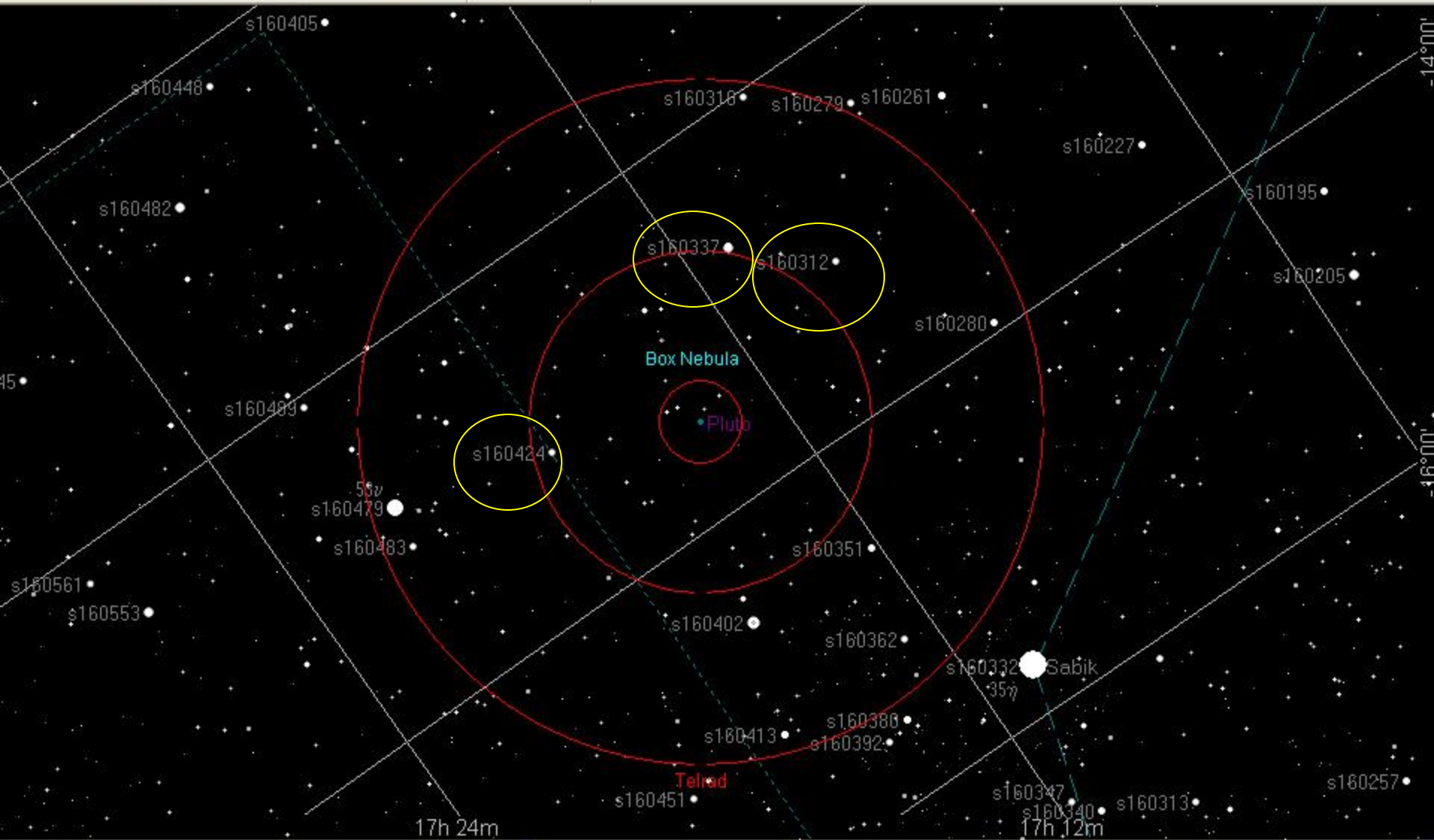
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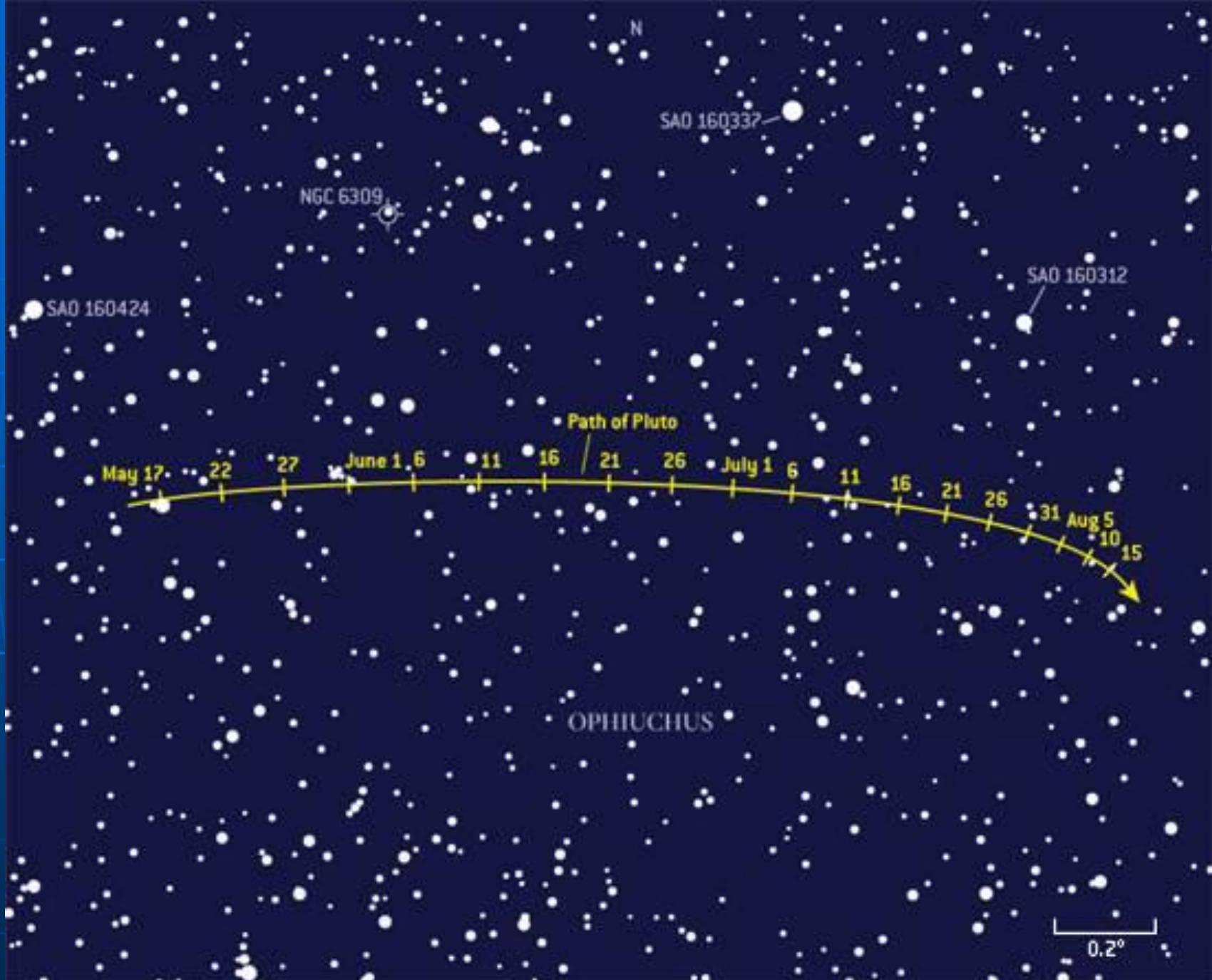




This wide-field view of the area in Ophiuchus where Pluto is located will help you get your bearings. The images following expand on the area contained in the green box.







S

OPHIUCHUS



SAO 160312



0.2°

SAO 160424

NGC 6309

SAO 160337

N

SAO 160337

NGC 6309

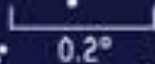
SAO 160312

SAO 160424

Path of Pluto

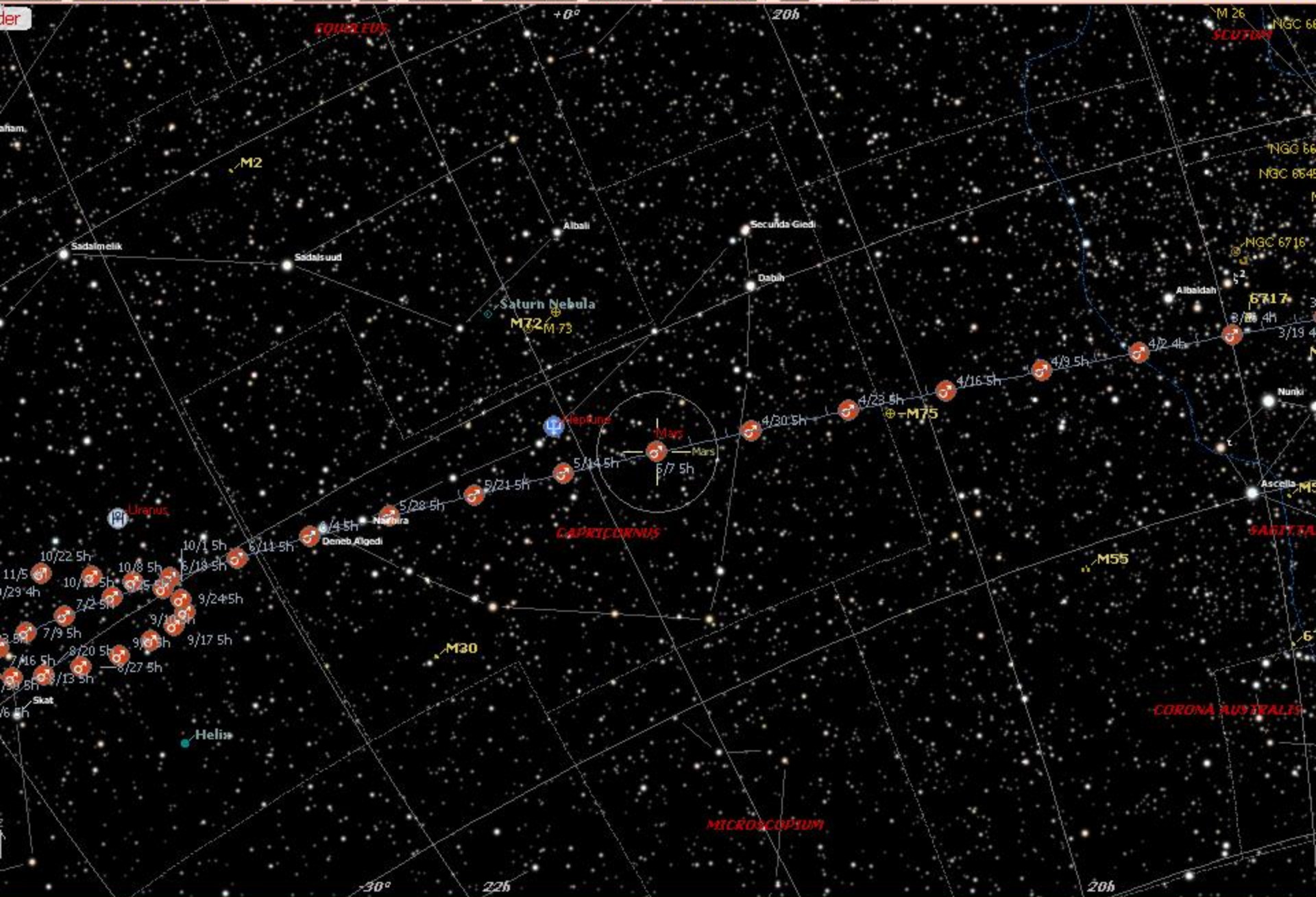


OPHIUCHUS



Mars:

Heading for Retrograde





The Projects for the Inner Solar System

MARS: Retrograde Motion

Early naked eye observers had a problem. The planet Mars, slowly drifting west to east from night to night, when seen against the background stars, would once a year act very strangely. As Mars approached opposition it would suddenly slow down, reverse itself, drift westward for a while (**retrograde motion**), before again reversing to assume its normal (**prograde**) eastward motion.

We now know that this is an illusion caused by the motion of the Earth catching up to and passing the slower Red Planet, causing Mars to appear to be moving backward.

You are to plot the apparent motion of Mars through this **retrograde** loop.

Determine what constellation Mars will be in at the time of opposition. This can be done by consulting the astronomy periodicals. Make a copy of that area out of a star atlas. For example, Will Tirion's Star Atlas 2000.0.

Then watch Mars beginning about a month before opposition until a month after opposition. Plot the planet's daily position on your copy by comparing its position to the fixed stars of the constellation.

After these two months you should be able to trace out Mars' **retrograde** motion.

Fortunately for us, the Copernican Revolution solved nicely the odd behavior of Mars, and also the behavior of Jupiter and Saturn, the other classical outer planets which exhibit a lesser amount of retrograde motion.